

2. Time evolution (Quantum dynamics)

2.1 Time evolution & the Schrödinger equation

Time in QM is a parameter ($|\psi(t)\rangle \in \mathcal{H}$),
not an observable like x .

Note: SR relates x, t ; restored in relativistic QFT, where
 x is no longer an observable.

Question: how does a state $|\psi(t)\rangle$ evolve in time?

Postulate (Schrödinger eq.)

$$i\hbar \frac{d}{dt} |\psi(t)\rangle = H |\psi(t)\rangle$$

Write
 In terms of time-evolution operator $U(t, t_0)$;

If state at time t_0 , $|\alpha, t_0\rangle \in \mathcal{H}$
 becomes at time t $|\alpha, t_0; t\rangle \in \mathcal{H}$,

$$\text{write } |\alpha, t_0; t\rangle = U(t, t_0) |\alpha, t_0\rangle.$$

Properties of $U(t, t_0)$:

i) Unitary - conserves probability, norm

$$\boxed{U^\dagger(t, t_0) U(t, t_0) = \mathbb{1}}$$

$$\begin{aligned} \langle \alpha, t_0; t | \alpha, t_0; t \rangle &= \langle \alpha, t_0 | U^\dagger(t, t_0) U(t, t_0) | \alpha, t_0 \rangle \\ &= \langle \alpha, t_0 | \alpha, t_0 \rangle. \end{aligned}$$

ii) composition law

$$\boxed{U(t, t_1) U(t_1, t_0) = U(t, t_0)}$$

$$\begin{aligned} |\alpha, t_0; t\rangle &= U(t, t_1) |\alpha, t_0; t_1\rangle \\ &= U(t, t_1) U(t_1, t_0) |\alpha, t_0\rangle \\ &= U(t, t_0) |\alpha, t_0\rangle \end{aligned}$$

iii) identity at $t = t_0$

$$\boxed{\lim_{t \rightarrow t_0} U(t, t_0) = \mathbb{1}} \quad \text{since } \lim_{t \rightarrow t_0} |\alpha, t_0; t\rangle = |\alpha, t_0\rangle.$$

Properties i) - iii) satisfied when infinitesimal form is

$$U(t_0 + dt, t_0) = 1 - \frac{iH(t_0)dt}{\hbar}$$

(equivalent to Schrödinger.)

Appearance of \hbar - needed on dimensional grounds.
- discuss further in ^{context of} classical-quantum correspondence

Schrödinger $\Leftarrow U(t, t_0)$

$$\boxed{i\hbar \frac{\partial}{\partial t} U(t, t_0) = H(t) U(t, t_0)} \quad (*)$$

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} |\alpha, t_0; t\rangle &= i\hbar \frac{\partial}{\partial t} U(t, t_0) |\alpha, t_0\rangle \\ &= H(t) U(t, t_0) |\alpha, t_0\rangle \\ &= H(t) |\alpha, t_0; t\rangle \end{aligned}$$

Solutions of (*).

1) Time-independent $H(t) = H$

$$\lim_{N \rightarrow \infty} \left[1 - \frac{i}{\hbar} H \frac{(t-t_0)}{N} \right]^N = e^{-\frac{iH}{\hbar}(t-t_0)}$$

so

$$U(t, t_0) = e^{-\frac{i}{\hbar} H (t-t_0)}$$

(can easily verify solves $i\hbar \frac{\partial}{\partial t} U(t, t_0) = H U(t, t_0)$).

2) Time-dependent, but $[H(t), H(t')] = 0$.

(Ex: particle in magnetic field, constant direction, varying strength)
 $H = \frac{p^2}{2m} + B(t) S_z$

similar solution but now

$$U(t, t_0) = e^{-\frac{i}{\hbar} \int_{t_0}^t H(t') dt'}$$

verify: $i\hbar \frac{\partial}{\partial t} U(t, t_0) = \frac{d}{dt} \left[\int_{t_0}^t H(t') dt' \right] U(t, t_0)$
 $= H(t) U(t, t_0)$

3) Time-dependent $H(t)$, $[H(t), H(t')] \neq 0$.

(Ex: particle in B field, direction changes in time.)

Solve iteratively

$$\int_{t_0}^t dt' \left[\frac{d}{dt'} U(t', t_0) \right] = -\frac{i}{\hbar} \int_{t_0}^t dt' H(t') U(t', t_0)$$

$$\Rightarrow U(t, t_0) = 1 - \frac{i}{\hbar} \int_{t_0}^t dt' H(t') U(t', t_0)$$

defines $U(t, t_0)$ in terms of $U(t', t_0)$, $t' \leq t$.

iterating:

$$U(t, t_0) = 1 - \frac{i}{\hbar} \int_{t_0}^t dt' H(t') \\ + \left(-\frac{i}{\hbar}\right)^2 \int_{t_0}^t dt'' \int_{t_0}^{t''} dt^{(2)} H(t^{(2)}) H(t^{(1)}) U(t^{(2)}, t_0)$$

$$= 1 + \sum_{n=1}^{\infty} \left(-\frac{i}{\hbar}\right)^n \int_{t_0}^t dt_1 \dots \int_{t_0}^{t_{n-1}} dt_n H(t_1) \dots H(t_n)$$

$$= \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{i}{\hbar}\right)^n \mathcal{T} \left(\int_{t_0}^t dt_1 \dots \int_{t_0}^{t_{n-1}} dt_n H(t_1) H(t_2) \dots H(t_n) \right)$$

(Dyson Series)

where \mathcal{T} is time-ordering operator - orders following ops so time goes up to left.

can write answer in compact form

$$U(t, t_0) = \mathcal{T} \left[e^{-\frac{i}{\hbar} \int_{t_0}^t dt' H(t')} \right]$$

(looks same as (2), but \mathcal{T} carries extra info above)

Time evolution of energy eigenkets (assume are (1); H is t -indep.)

Assume $\{|a\rangle\}$ is a complete basis of kets so that

$$H|a\rangle = E_a|a\rangle.$$

The time-evolution operator $U(t, t_0)$ is

$$U(t, t_0) = e^{-\frac{i}{\hbar} H(t-t_0)} = \sum_a |a\rangle e^{-\frac{i}{\hbar} E_a(t-t_0)} \langle a|.$$

$$\text{If } |\alpha, t_0\rangle = \sum C_a(t_0) |a\rangle,$$

$$\text{then } |\alpha, t_0=0; t\rangle = \sum C_a(t) |a\rangle$$

$$\text{where } C_a(t) = e^{-\frac{i}{\hbar} E_a t} C_a(0).$$

Note: only phases change under time-development,
probability $|C_a(t)|^2$ of being in state $|a\rangle$
is unchanged.

Useful to find CSCO A_1, \dots, A_k so that
 $[A_1, H] = [A_2, H] = \dots = [A_k, H] = 0$

so can find a basis $|a_1, \dots, a_k\rangle$ of H eigenkets.

2.2 Schrödinger, Heisenberg, & interaction pictures

Previous discussion used Schrödinger picture:

$|\alpha, t\rangle \in \mathcal{H}$ evolves in time, operators fixed.

Two ways to view expectation values:

$$\begin{array}{c} \text{Schrödinger:} \\ \langle A \rangle = \underbrace{\langle \alpha, t |}_{(S)} \underbrace{A}_{(S)} \underbrace{|\alpha, t\rangle}_{(S)} \\ \langle A \rangle = \underbrace{\langle \alpha, 0 |}_{(H)} \underbrace{U^\dagger(t, 0) A U(t, 0)}_{(H)} \underbrace{|\alpha, 0\rangle}_{(H)} \\ \text{Heisenberg} \end{array}$$

Heisenberg picture: operators evolve in time, state fixed

Same physics - different formalism.

Convention: set equal at $t=0$

$$|\alpha, 0\rangle_{(S)} = |\alpha\rangle_{(H)}$$

$$A_{(H)}(0) = A_{(S)}$$

then

$$|\alpha, t\rangle_{(S)} = U(t, 0) |\alpha\rangle_{(H)} \quad A_{(H)}(t) = U^\dagger(t, 0) A_{(S)} U(t, 0)$$

If H time-independent,

$$|\alpha, t\rangle_{(S)} = e^{-\frac{i}{\hbar} H t} |\alpha\rangle_{(H)} \quad A_{(H)}(t) = e^{\frac{i}{\hbar} H t} A_{(S)} e^{-\frac{i}{\hbar} H t}$$

Heisenberg equation of motion ~~(A possibly t-independent)~~ (A possibly t-dependent)

$$\begin{aligned} \frac{d}{dt} A_{(H)}(t) &= \frac{\partial U^\dagger}{\partial t} A_{(S)} U + U^\dagger A_{(S)} \frac{\partial U}{\partial t} + U^\dagger \frac{\partial A_{(S)}}{\partial t} U \\ &= \frac{i}{\hbar} U^\dagger H \underbrace{(U U^\dagger)}_1 A_{(S)} U - \frac{i}{\hbar} U^\dagger A_{(S)} (U U^\dagger) H U + U^\dagger \frac{\partial A_{(S)}}{\partial t} U \end{aligned}$$

if case (1) or (2), $U^\dagger H U = H$, so $H_{(H)} = H$,

$$\boxed{\frac{d}{dt} A_{(H)}(t) = \frac{i}{\hbar} [A_{(H)}(t), H] + \dot{A}_{(H)}}_{\substack{\uparrow \\ \text{vanishes if } A_{(S)} \\ \text{is } t\text{-independent.}}}$$

Interaction picture

Sometimes useful to use a "split picture"

$$\text{Consider } H = \underbrace{H_0}_{\text{time-independent}} + \underbrace{V(t)}_{\text{time dependent}}$$

Interaction picture: remove H_0 evolution from state, as in Heisenberg.

$$|\alpha\rangle_{(H)} = e^{\frac{i}{\hbar} H t} |\alpha, t\rangle_{(S)}$$

$$|\alpha\rangle_{(I)} = e^{\frac{i}{\hbar} H_0 t} |\alpha, t\rangle_{(S)}$$

$$A_{(H)} = e^{\frac{i}{\hbar} H t} A_{(S)} e^{-\frac{i}{\hbar} H t}$$

$$A_{(I)} = e^{\frac{i}{\hbar} H_0 t} A_{(S)} e^{-\frac{i}{\hbar} H_0 t}$$

Equation of motion in interaction picture

$$\begin{aligned}
 i\hbar \frac{\partial}{\partial t} |\alpha, t\rangle_{(I)} &= -H_0 e^{\frac{i}{\hbar} H_0 t} |\alpha, t\rangle_{(S)} + e^{\frac{i}{\hbar} H_0 t} (H_0 + V) |\alpha, t\rangle_{(S)} \\
 &= \underbrace{e^{\frac{i}{\hbar} H_0 t} V e^{-\frac{i}{\hbar} H_0 t}}_{V_I} \underbrace{e^{\frac{i}{\hbar} H_0 t} |\alpha, t\rangle_{(S)}}_{|\alpha, t\rangle_{(I)}} \\
 &= V_I |\alpha, t\rangle_{(I)}.
 \end{aligned}$$

s.
$$i\hbar \frac{\partial}{\partial t} |\alpha, t\rangle_{(I)} = V_I |\alpha, t\rangle_{(I)}$$
 evolves with V .

$$\frac{dA_{(I)}}{dt} = \frac{1}{i\hbar} [A_{(I)}, H_0] + \dot{A}_{(I)}$$

evolves with H .

Summary:

	State	Operator
Schrödinger	evolves w/ H	const.
Heisenberg	constant	evolves w/ H
Interaction	evolves w/ V_I	evolves w/ H_0

Will return to this picture for time-independent pert. thry.

Base kets & transition amplitudesSchrödinger: State ket $|\psi(t)\rangle$ changesHeisenberg: " " $|\psi\rangle$ doesn't change.

Schrödinger eqn: $i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle_{\text{S}} = H |\psi(t)\rangle_{\text{S}}$

Heisenberg eqn: $\frac{dA_H(t)}{dt} = \frac{1}{i\hbar} [A_H(t), H] + \dot{A}_H(t)$

$|\psi(t)\rangle = U(t, t_0) |\psi(t_0)\rangle$

$A_H(t) = U^\dagger(t, t_0) A_{\text{int}}(t_0) U(t, t_0)$

Given a (time independent) operator A ,
($\dot{A} = 0$)in Schrödinger picture ~~the~~ states $|a'\rangle$ satisfy

$$A |a'\rangle = a' |a'\rangle$$

don't change in time.

Heisenberg:

$$A_H(t) = U^\dagger A(0) U$$

$$A_H(U^\dagger |a'\rangle) = U^\dagger A(0) |a'\rangle = a' (U^\dagger |a'\rangle)$$

so $|a', t\rangle_{\text{H}} = U^\dagger |a'\rangle$

Base kets change in time in H. picture (Eigenvalues unchanged)

Two interpretations:

$$C_{a'} = \underbrace{\langle a' |}_{\text{S: base}} \underbrace{U}_{\text{H: base}} \underbrace{|a, t=0\rangle}_{\text{state}}$$

Transition amplitude if $A = a'$ at time $t = 0$, what is prob. $B = b'$ at time t

$$\underbrace{\langle b' |}_{\text{S: base}} \underbrace{U(t, 0)}_{\text{H: base}} \underbrace{|a'\rangle}_{\text{state}}$$

Energy - time uncertainty relation

Unlike x , t is not an operator, so no direct analog of $\Delta x \Delta p \geq \hbar/2$ ($\langle \Delta x^2 \rangle \langle \Delta p^2 \rangle \geq \hbar^2/4$)

Q: how rapidly does a state change form?

Define $C(t) = \langle \alpha | U(t, t_0) | \alpha \rangle$

(Don't confuse w/ $C(t)$ from prob. 15 in bk.)

If $|\alpha\rangle$ an eigenvector of H , $|C(t)| = 1$, $\forall t$.
("stationary state")

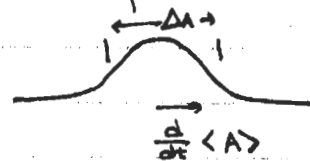
Generally, $|\alpha\rangle = \sum c_\alpha |\alpha\rangle$

$$C(t) = \sum |c_\alpha|^2 e^{-\frac{iE_\alpha t}{\hbar}}$$

as t increases, generically $|C(t)|$ decreases.

Imagine measuring an observable A which changes in time
- use a clock (i.e., position of particle, hands of clock, ...)

Can measure $\Delta E = \frac{\Delta A}{\frac{d}{dt} \langle A \rangle}$



$$\frac{d}{dt} \langle A \rangle = \frac{1}{i\hbar} \langle [A, H] \rangle$$

$$\langle \Delta A^2 \rangle \langle \Delta H^2 \rangle \geq \frac{1}{4} |\langle [A, H] \rangle|^2 = \frac{\hbar^2}{4} \left| \frac{d}{dt} \langle A \rangle \right|^2$$

$$\text{So } \frac{\langle \Delta A^2 \rangle}{\left| \frac{d}{dt} \langle A \rangle \right|^2} \langle \Delta H^2 \rangle \geq \hbar^2/4$$

$$\boxed{\Delta T \Delta E \geq \hbar/2}$$

$$\Delta E = \langle \Delta H^2 \rangle^{1/2}$$

$$\Delta T = \left(\frac{\langle \Delta A^2 \rangle}{\left| \frac{d}{dt} \langle A \rangle \right|^2} \right)^{1/2}$$

Basic idea: if energy width is small, ^{forming state} ~~particle~~ takes a long time to change.

Interpretation of wavefunction ("probability fluid")

Start with Schrödinger picture for particle in 3D potential $i\hbar \frac{\partial}{\partial t} \psi(\vec{x}, t) = H \psi(\vec{x}, t) = -\frac{\hbar^2}{2m} \nabla^2 \psi(\vec{x}, t) + V(\vec{x}) \psi(\vec{x}, t)$

Think about $\rho(\vec{x}) = |\psi(\vec{x}, t)|^2$ as probability density

$$[\text{probability} (\text{at } \vec{x} \in \mathcal{R})] = \int_{\mathcal{R}} |\psi(\vec{x}', t)|^2 d^3x' \quad \text{[R]}$$

Compute $\frac{\partial \rho}{\partial t}$ for $\rho(\vec{x}, t)$ in 3D

$$\frac{\partial \rho}{\partial t} = \psi^* \frac{\partial \psi}{\partial t} + \frac{\partial \psi^*}{\partial t} \psi$$

$$\frac{\partial \psi}{\partial t} = i\frac{\hbar}{2m} \nabla^2 \psi - \frac{iV}{\hbar} \psi$$

$$\frac{\partial \psi^*}{\partial t} = -i\frac{\hbar}{2m} \nabla^2 \psi^* + \frac{iV}{\hbar} \psi^*$$

cancel, since V real

$$\Rightarrow \frac{\partial \rho}{\partial t} = -\frac{i\hbar}{2m} [(\nabla^2 \psi^*) \psi - \psi^* (\nabla^2 \psi)]$$

$$= -\frac{i\hbar}{2m} \vec{\nabla} \cdot [(\vec{\nabla} \psi^*) \psi - \psi^* (\vec{\nabla} \psi)]$$

$$= -\vec{\nabla} \cdot \left[\frac{\hbar}{m} \text{Im}(\psi^* \vec{\nabla} \psi) \right]$$

$\vec{j}(\vec{x}, t)$ "probability flux"

$$\text{so } \boxed{\frac{\partial \rho}{\partial t} = -\vec{\nabla} \cdot \vec{j}(\vec{x}, t)}$$

continuity equation

\vec{j} has natural interpretation as flux vector for probability.

$$\left(\frac{d}{dt} \int_V \rho dV = - \int_{\partial V} \vec{j} \cdot d\vec{A} \right)$$

\vec{j} related to momentum

$$\int d^3\vec{x} \, j(\vec{x}, t) = \frac{1}{m} \int \psi^*(\vec{x}, t) (-i\hbar \vec{\nabla}) \psi(\vec{x}, t)$$

$$= \frac{1}{m} \langle \psi(t) | \vec{p} | \psi(t) \rangle$$

Physical significance of phase

write $\psi(\vec{x}, t) = \sqrt{\rho(\vec{x}, t)} e^{\frac{iS(\vec{x}, t)}{\hbar}}$

↑
amplitude

↑
phase

$$\psi^* \vec{\nabla} \psi = \frac{1}{2} \vec{\nabla} \rho + \frac{i}{\hbar} \rho \vec{\nabla} S$$

$$\therefore \vec{j}(\vec{x}, t) = \frac{1}{m} \rho(\vec{x}, t) \vec{\nabla} S(\vec{x}, t)$$

So: rate of variation of S controls flow of probability.
Faster phase variation \rightarrow more prob. flow

Ex. stationary bound state: $\psi(\vec{x}, t)$ has constant phase
(can choose real @ $t=0$)
 \rightarrow no flow of probability

Ex. Plane wave $\psi(\vec{x}, t) \approx e^{\frac{i p x}{\hbar} - \frac{i E t}{\hbar}}$

$$\vec{\nabla} S = \vec{p}$$

So $\frac{1}{m} \vec{\nabla} S$ is like velocity " \vec{v} "

$$\frac{\partial \rho}{\partial t} \approx \vec{\nabla}(\rho \cdot \vec{v}).$$

Suggestive, like fluid mechanics. Gives intuition, but not to be taken literally.

2.3 Connections between Classical & Quantum Mechanics

Review of Classical physics

3 Approaches:

A) Newton

$$\text{EOM: } F = ma$$

Ex. 1D SHO with potential $V(x) = \frac{1}{2}m\omega^2 x^2$

$$m\ddot{x} = -\frac{d}{dx}V(x) = -m\omega^2 x \quad [= -kx, \omega = \sqrt{k/m}]$$

B) Hamiltonian

Phase space (x 's & p 's) with Poisson bracket

$$\{X^i, P^j\} = \delta^{ij} \quad (\text{locally})$$

Ham. function H

$$\text{EOM: } \dot{q} = \{q, H\}$$

$$\text{Ex. SHO} \quad H = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2$$

$$\dot{x} = \{x, H\} = p/m$$

$$\dot{p} = \{p, H\} = -m\omega^2 x$$

C) Lagrangian (principle of least action)

Start with Lagrangian $\mathcal{L}(x^i, \dot{x}^i)$

[Related to Hamiltonian through $H = p_i \dot{x}^i - \mathcal{L}$]

Define Action $S[x(t)]$ as functional on space of paths

$$S = \int dt \mathcal{L}(x^i, \dot{x}^i)$$

Classical trajectory extremizes S

$\delta S = 0 \Rightarrow$ Euler-Lagrange equations

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{x}^i} - \frac{\partial \mathcal{L}}{\partial x^i} = 0$$

Ex. SHO

$$\mathcal{L} = \frac{1}{2} m \dot{x}^2 - \frac{1}{2} m \omega^2 x^2$$

$$\frac{d}{dt} m \dot{x} + m \omega^2 x = 0$$

$$\Rightarrow m \ddot{x} = -m \omega^2 x$$

S Related to Hamilton's principle function (or in WKB)
through

$$S[x, t; x_0, t_0] = S[x_{\text{class}}(t)]$$

$$= \int_{t_0}^t dt \mathcal{L}(x^i, \dot{x}^i) \quad \text{along classical trajectory}$$

from $t_0, x_0 \rightarrow x, \dots$

Relating Classical & Quantum mechanics

A) Ehrenfest

Consider a particle in a 3D potential $V(\vec{x})$

$$H = \frac{\vec{p}^2}{2m} + V(\vec{x})$$

$$= -\frac{\hbar^2}{2m} \nabla^2 + V(\vec{x})$$

Use Heisenberg equation to write $\langle \frac{d\vec{x}}{dt} \rangle$, $\langle \frac{d^2\vec{x}}{dt^2} \rangle$

$$\frac{d\vec{x}}{dt} = \frac{1}{i\hbar} [\vec{x}, H] = \frac{\vec{p}}{m}$$

$$\text{so } \langle \frac{d\vec{x}}{dt} \rangle = \frac{1}{m} \langle \vec{p} \rangle$$

$$\frac{d^2\vec{x}}{dt^2} = \frac{1}{i\hbar} \left[\frac{\vec{p}}{m}, V(x) \right] = -\frac{1}{m} \nabla V(x).$$

$$\text{So } \boxed{m \frac{d^2}{dt^2} \langle \vec{x} \rangle = \frac{d}{dt} \langle \vec{p} \rangle = -\langle \nabla V(x) \rangle}$$

Ehrenfest's theorem

Classical EOM emerges - note: no \hbar !

Generally, for any system described by classical physics, classical description can be derived from QM starting point.

Not all systems have classical limits (eg. 2-state system)

B) Quantization & Hamiltonian Mechanics

In principle, all quantum systems (not including gravity) described by Standard model (Quantum field theory)

Sometimes we want to "guess" underlying quantum system, given classical description: Quantization

Often, can be done by taking $\{ \cdot, \cdot \} \rightarrow [\cdot, \cdot]$
through

$$\{f, g\} = h \Rightarrow [F, G] = i\hbar H$$

For example, $\{x, p\} = 1 \rightarrow [X, P] = i\hbar \mathbb{1}$.

This program can encounter ambiguities due to Operator ordering problems.

Ex: $xp \neq px$, so how to quantize xp operator?

Can use Hermiticity as guideline

$\rightarrow \frac{1}{2}(xp + px)$ is Hermitian.

But this doesn't always work. Generally, need to try various possibilities.

$$[\text{Ex. } xp^2 \rightarrow \frac{1}{2}[x^2p^2 + p^2x^2] = xp^2x - \hbar^2]$$

This trial & error process led to many current QM models.

Quantization takes Hamiltonian EOM

$$\frac{d}{dt} q = \{q, H\}$$

to Heisenberg EOM $i\hbar \frac{d}{dt} A = [A, H]$

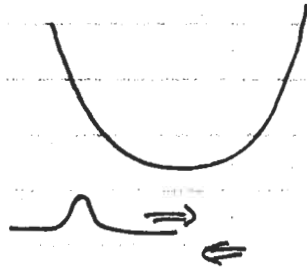
(This is why Ehrenfest works.)

Classical picture emerges from QM in limit $\hbar \rightarrow 0$.

Wavefunction "close to" eigenstate of all relevant classical operators



Particularly nice example: coherent states of SHO.
Retain shape, act like classical states



slush back & forth.

[\Rightarrow CD WKB]

C2) Propagators & path integrals

Recall time-development

$$|\psi_a(t)\rangle = \sum_{a'} C_{a'}(t) |a'\rangle$$

$$C_{a'}(t) = e^{-\frac{i}{\hbar} E_{a'}(t-t_0)} C_{a'}(t_0)$$

For particle in 1D/3D

If ~~state~~ $\langle x|a'\rangle = U_{a'}(x)$,

$$\psi(x,t) = \sum_{a'} e^{-\frac{i}{\hbar} E_{a'}(t-t_0)} C_{a'}(t_0) U_{a'}(x).$$

c) WKB approximation

Quasi-classical approximation

$$\psi = \sqrt{\rho} e^{iS/\hbar}$$

Expand $i\hbar \frac{\partial}{\partial t} \psi = -\frac{\hbar^2}{2m} \nabla^2 \psi + V(x) \psi = E\psi$ in \hbar .

$\mathcal{O}(\hbar)$ terms:

$$-\frac{\partial S}{\partial t} = V + \frac{1}{2m} |\nabla S|^2$$

(Hamilton-Jacobi eqn: satisfied by H. principal fn.)

Look at a stationary state in 1D

$$\frac{1}{2m} (S')^2 = E - V$$

$$\Rightarrow S(x) = \pm \int \sqrt{2m(E-V)} dx'$$

$$= \pm \int p dx \quad p = \sqrt{2m(E-V)}$$

$\mathcal{O}(\hbar^2)$ terms:

$$\frac{\partial \rho}{\partial t} = -\frac{1}{m} \frac{\partial}{\partial x} \left(\rho \frac{\partial S}{\partial x} \right) = 0 \quad (\text{continuity eqn})$$

$$= -\frac{1}{m} \frac{\partial}{\partial x} \left[\rho \sqrt{2m(E-V)} \right]$$

$$\Rightarrow \rho = \frac{\text{const}}{\sqrt{2m(E-V)}} = \frac{C}{\sqrt{p}}$$

(physical interp: time spent in region of mom. p
 $\sim 1/p$ - agrees w/ classical intuition)

So for a stationary bound state

$$\psi(x) = \frac{c_1}{\sqrt{p}} e^{\frac{i}{\hbar} \int p dx} + \frac{c_2}{\sqrt{p}} e^{-\frac{i}{\hbar} \int p dx}$$

$$p = \sqrt{2m(E-V)}$$

This is WKB approximation

valid when $\hbar S'' \ll (S')^2$

$$\Leftrightarrow \left| \frac{d}{dx} \left(\frac{\hbar}{S'} \right) \right| \ll 1$$

$$\frac{d}{dx} \left(\frac{\hbar}{\sqrt{2m(E-V)}} \right) = \frac{2m\hbar V'(x)}{2(2m(E-V))^{3/2}}$$

so condition for validity is

$$\lambda = \frac{\hbar}{p} \ll \frac{2(E-V)}{V'}$$

\nearrow distance over which V changes appreciably.

WKB valid in short wavelength limit,

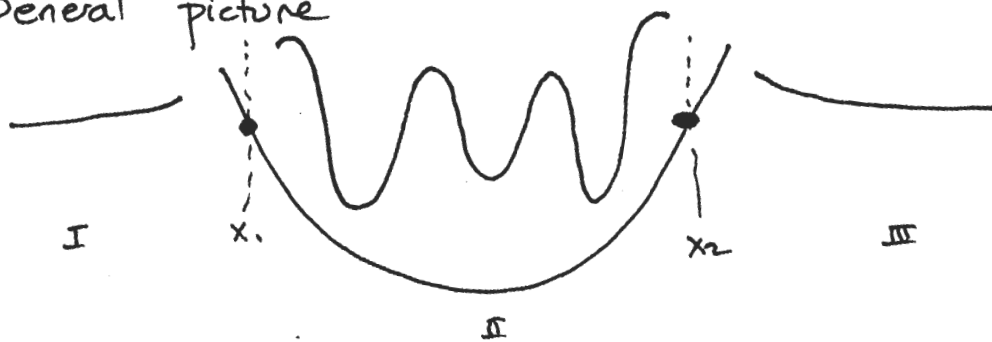
not near $E = V(x)$ (classical turning points)

Still valid when $E < V(x)$ though

$$\psi(x) = \frac{c_{\pm}}{\sqrt{2m(V-E)}} e^{\pm \frac{i}{\hbar} \int \sqrt{2m(V-E)} dx}$$

[Only one term is valid - take exponential damping for bd. states]

General picture



know ψ in regions I, II, III,
must match behavior @ x_1, x_2

(like exact solution in \square potential)
or \square

Region II:
$$\psi = \frac{C_1}{p^{1/2}} e^{i \int_{x_2}^x p dx'} + \frac{C_2}{p^{1/2}} e^{-i \int_{x_2}^x p dx'}$$

III:
$$\psi = \frac{C}{|p|^{1/2}} e^{-\frac{1}{\hbar} \int_{x_2}^x |p| dx'}$$

One approach: use exact solution near x_2 : $V(x) \sim E + F(x-x_2)$

Airy functions
$$\mathbb{F}(u) \sim \int_0^{\infty} \cos(ux + \frac{1}{3}u^3) du \sim \begin{cases} J_{2/3}(\frac{2}{3}|x|^{3/2}) & \text{II} \\ K_{1/3}(\frac{2}{3}|x|^{3/2}) & \text{III} \end{cases}$$

using asymptotic behaviour, match to ψ in regions II, III

Clearer approach: analytic continuation in x plane, away from x

$$\begin{aligned} \psi^{(III)} &= \frac{C}{\sqrt{2mF(x-x_2)}} e^{-\frac{1}{\hbar} \int_{x_2}^x \sqrt{2mF(x'-x_2)} dx'} \\ &= \frac{C}{\sqrt{2mF(x-x_2)}} e^{-\frac{2}{3\hbar} \sqrt{2mF} (x-x_2)^{3/2}} \end{aligned}$$

say $x = x_2 + \rho e^{i\phi} \Rightarrow (x-x_2)^{3/2} = \rho^{3/2} e^{\frac{3}{2}i\phi}$

if $x = x_2 - \rho$, take $\phi = \pi$, $(x-x_2)^{3/2} = \rho^{3/2} (-i)$

so $\psi^{(III)} \rightarrow \frac{C e^{-i\pi/4}}{(2mF(x_2-x))^{1/4}} e^{\frac{2i}{3\hbar} \sqrt{2mF} (x_2-x)^{3/2}}$

(analytic cont. in UHP \curvearrowright)

matches with C_2 term in $\psi^{(II)}$,

$$C_2 = C e^{i\pi/4}$$

similarly $C_1 = C e^{i\pi/4}$ (analytic cont. in LHP \curvearrowright)

so $\psi^{(I)} = \frac{C}{(2mF(x_2-x))^{1/4}} \cos \left[-\frac{1}{\hbar} \int_x^{x_2} \sqrt{2mF(x_2-x')} dx' + \frac{\pi}{4} \right]$

when

$$\psi^{(III)} = \frac{C}{(2mF(x-x_2))^{1/4}} e^{-\frac{1}{\hbar} \int_{x_2}^x \sqrt{2mF(x'-x_2)} dx'}$$

Using I/II & I/II overlaps.

$$\psi^{\text{inside}} = \frac{C}{(E-V)^{1/4}} \cos \left[-\frac{1}{\hbar} \int_x^{x_2} \sqrt{2m(E-V(x'))} dx' + \frac{\pi}{4} \right]$$

$$= \frac{C}{(E-V)^{1/4}} \cos \left[\frac{1}{\hbar} \int_{x_1}^x \sqrt{2m(E-V(x'))} dx' - \frac{\pi}{4} \right]$$

but wavefunction is unique, so

$$\int_{x_1}^{x_2} dx' \sqrt{2m(E-V(x'))} = (n + \frac{1}{2}) \pi \hbar$$

[like Bohr-Sommerfeld ~~except~~ $\frac{1}{2}$]

WKB approximation for bound state energies.

Improves as $n \rightarrow \infty$, since $\lambda \rightarrow 0$

SHO: derive K in homework.

Properties of K :

Quantum stat. mech.

$$\text{Define } G(t) = \int d^3x K(x, t; x', t_0) \\ = \sum_{a'} e^{-\frac{iE_a t}{\hbar}}$$

$$\text{set } t = -i\hbar\beta,$$

$$G(-i\hbar\beta) = Z = \sum_{a'} e^{-\beta E_{a'}}$$

(related to QMC)

stat. mech. partition function $\beta \sim \frac{1}{T}$

Fourier transform.

$$\text{Define } \tilde{G}(E) = -i \int dt G(t) e^{iEt} \\ = -i \sum_{a'} \int_0^{\infty} dt e^{i(E - E_{a'})t}$$

For convergence, take $E + i\epsilon$

$$\tilde{G}(E + i\epsilon) = \sum_{a'} \frac{\hbar}{E - E_{a'} + i\epsilon}$$

poles in limit $\epsilon \rightarrow 0$ describe energy spectrum.

Density of states

$$\rho(E) = \sum_{\alpha} \delta(E - E_{\alpha}) \quad \text{for discrete spectrum.}$$

$$\pi \delta(E - E') = \lim_{\epsilon \rightarrow 0} \frac{\epsilon}{(E - E')^2 + \epsilon^2} = \lim_{\epsilon \rightarrow 0} \text{Im} \frac{1}{E - E' + i\epsilon}$$

$$\text{so } \rho_{\epsilon}(E) = \frac{1}{\pi \hbar} \text{Im} \tilde{G}(E + i\epsilon)$$

is regulated state density.

Path integrals

Note composition property of K :

$$K(x, t; x', t_0) = \int d\tilde{x} K(x, t; \tilde{x}, \tilde{t}) K(\tilde{x}, \tilde{t}; x', t_0)$$

[valid to Kret for $t_0 < \tilde{t} < t$]

$$\text{(follow from } U(t, \tilde{t}) U(\tilde{t}, t_0) = U(t, t_0)\text{)}$$

Break $t - t_0$ into N equal time intervals

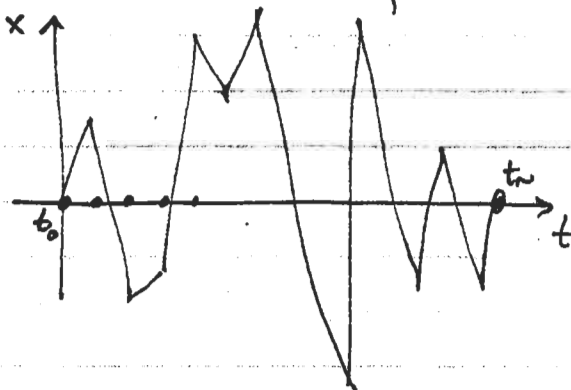
$$\Delta t = \frac{t - t_0}{N}$$

$$t_k = t_0 + k \Delta t$$

$$t_N = t$$

$$\text{then } K(x_N, t_N; x_0, t_0) = \int \prod_{k=1}^{N-1} dx_k K(x_N, t_N; x_{N-1}, t_{N-1}) \\ \cdot K(x_{N-1}, t_{N-1}; x_{N-2}, t_{N-2}) \\ \dots \cdot K(x_1, t_1; x_0, t_0)$$

so final answer includes all paths



Feynman proposed:

$$K(x'', t, x', t_0) = \int \mathcal{D}[x(t)] e^{\frac{i}{\hbar} S[x(t)]}$$

where $\mathcal{D}[x(t)]$ is a measure on the space of paths with $x(t_0) = x'$, $x(t) = x''$.

- Clearly obeys composition rule
- Simple connection to classical physics - phases cancel except near stationary point $\delta S = 0$.

To make rigorous, must define measure on path space
[Wiener measure, etc... [now used in economics (finance, etc.)]]

Plan: start from definition of K .

"Derive" PI & appropriate measure,
go back & rederive K for free particle

$$\begin{aligned}
 K(X_N, t_N; X_0, t_0) &= \int \prod_{k=1}^{N-1} dX_k \langle X_N | U(t_N, t_{N-1}) | X_{N-1} \rangle \\
 &\quad \langle X_{N-1} | U(t_{N-1}, t_{N-2}) | X_{N-2} \rangle \\
 &\quad \dots \langle X_1 | U(t_1, t_0) | X_0 \rangle \\
 &= \int \prod_{k=1}^{N-1} dX_k \langle X_N | e^{-\frac{i\epsilon}{\hbar} H} | X_{N-1} \rangle \langle X_{N-1} | e^{-\frac{i\epsilon}{\hbar} H} | X_{N-2} \rangle \\
 &\quad \dots \langle X_1 | e^{-\frac{i\epsilon}{\hbar} H} | X_0 \rangle
 \end{aligned}$$

[$\epsilon = \Delta t$]

Note: easy to include t -dependent H , time ordering works out automatically but will ignore for clarity.

write

$$\begin{aligned}
 \langle X_k | e^{-\frac{i\epsilon}{\hbar} H(p, x)} | X_{k-1} \rangle \\
 = \int dp_k \langle X_k | p_k \rangle \langle p_k | e^{-\frac{i\epsilon}{\hbar} H(p, x)} | X_{k-1} \rangle
 \end{aligned}$$

Introduce notation: Normal ordering.

$\mathcal{O}(p, x)$ is a normal-ordered operator if p 's on left, x 's on right.
[often use NO notation for a^\dagger, a 's]

Ex. $H = \frac{p^2}{2m} + V(x)$ is normal ordered.

Write : $\mathcal{O}(p, x)$: for normal-ordered form of \mathcal{O} .

$$\text{Ex. } :xp: = \overbrace{px} + i\hbar = :xp: + i\hbar$$

[normal ordering introduces commutators]

$$\text{Ex: } H = \frac{1}{2m} (\vec{p} - \frac{e}{c} \vec{A}(\vec{x}))^2 \quad (\text{particle in EM field})$$

$$H = \frac{1}{2m} (\vec{p}^2 - \frac{e}{c} [\vec{p} \cdot \vec{A}(x) + \vec{A}(x) \cdot \vec{p}] + \frac{e^2}{c^2} A^2(x))$$

$$= :H: - \frac{ie}{c} \vec{\nabla} \cdot \vec{A}(x)$$

would like $e^{-\frac{i\epsilon}{\hbar} H(p,x)}$ to be normal ordered.

For $H = \frac{p^2}{2m} + V(x)$,

$$\begin{aligned}
 e^{-\frac{i\epsilon}{\hbar} H(p,x)} &= 1 - \frac{i\epsilon}{\hbar} \left[\frac{p^2}{2m} + V(x) \right] \\
 &\quad - \frac{\epsilon^2}{2\hbar^2} \left[\left(\frac{p^2}{2m} \right)^2 + \frac{p^2}{2m} V(x) + V(x) \frac{p^2}{2m} + V(x)^2 \right] \\
 &= : e^{-\frac{i\epsilon}{\hbar} H(p,x)} : \underbrace{- \frac{\epsilon^2}{2\hbar^2} \left[V(x), \frac{p^2}{2m} \right]}_{-\frac{\epsilon^2}{4m} \left[2\frac{i}{\hbar} V'(x)p - V''(x) \right]}
 \end{aligned}$$

Generally, if $H(p,x)$ is normal-ordered,

$$e^{-\frac{i\epsilon}{\hbar} H(p,x)} = : e^{-\frac{i\epsilon}{\hbar} H(p,x)} : + \mathcal{O}(\epsilon^2).$$

as $\Delta t \rightarrow 0$, replace $e^{-\frac{i\epsilon}{\hbar} H(p,x)} \rightarrow : e^{-\frac{i\epsilon}{\hbar} H(p,x)} :$

so $\int dp_k \langle x_k | p_k \rangle \langle p_k | e^{-\frac{i\epsilon}{\hbar} H(p,x)} | x_{k-1} \rangle$

becomes

$$\int dp_k \left(\frac{1}{2\pi\hbar} \right) e^{\frac{i}{\hbar} p_k (x_k - x_{k-1}) - \frac{i\epsilon}{\hbar} H(p_k, x_{k-1})} + \mathcal{O}(\epsilon^2)$$

so $K(x_N, t_N; x_0, t_0) \approx \int \left(\prod_{k=1}^{N-1} dx_k \right) \left(\prod_{k=1}^N \frac{dp_k}{2\pi\hbar} \right) e^{\sum_{k=1}^N \left[\frac{i}{\hbar} p_k (x_k - x_{k-1}) - \frac{i\epsilon}{\hbar} H(p_k, x_{k-1}) \right]}$

Replacing $\frac{x_k - x_{k-1}}{\epsilon} \rightarrow \dot{x}$

$$\sum_k \epsilon f_k \rightarrow \int dt f(t)$$

$$\left(\prod_{k=1}^{N-1} dx_k \right) \left(\prod_{k=1}^N \frac{dp_k}{2\pi\hbar} \right) \rightarrow \mathcal{D}[x(t)] \mathcal{D}[p(t)]$$

[Functional measure defined by limit]

gives phase space form of path integral:

$$K(x_N, t_N; x_0, t_0) = \int \mathcal{D}[x(t)] \mathcal{D}[p(t)] e^{\frac{i}{\hbar} \int dt [p(t) \dot{x}(t) - H(p(t), x(t))]}$$

Lagrangian form of PI

say $H = \frac{p^2}{2m} + V(x)$

$$\frac{1}{2\pi\hbar} \int dp_k e^{\frac{i}{\hbar} p_k (x_k - x_{k-1}) - \frac{i\epsilon}{\hbar} H(p_k, x_{k-1})}$$

$$= \frac{1}{2\pi\hbar} \int dp_k e^{\frac{-i\epsilon}{2m\hbar} \left[\left(p_k - \frac{m}{\epsilon} (x_k - x_{k-1}) \right)^2 - \frac{m^2}{\epsilon} (x_k - x_{k-1})^2 \right] - \frac{i\epsilon}{\hbar} V(x_{k-1})}$$

$$\int e^{-\frac{a}{2}x^2} = \sqrt{\frac{\pi}{a}}$$

$$= \sqrt{\frac{m}{2\pi i \hbar \epsilon}} e^{\frac{im^2}{2\hbar\epsilon} (x_k - x_{k-1})^2 - \frac{i\epsilon}{\hbar} V(x_{k-1})}$$

so

$$K(x, t; x_0, t_0) \approx \underbrace{\left(\frac{m}{2\pi i \hbar \epsilon} \right)^{N/2} \int \prod_{k=1}^{N-1} dx_k}_{\mathcal{D}[x(t)]} e^{\sum_{k=0}^{N-1} \left[-\frac{i\epsilon}{\hbar} V(x_{k+1}) + \frac{im^2}{2\hbar\epsilon} (x_k - x_{k+1})^2 \right]}$$

$$K(x, t; x_0, t_0) = \int \mathcal{D}[x(t)] e^{\frac{i}{\hbar} \int dt \left[\frac{1}{2} m \dot{x}(t)^2 - V(x) \right]}$$

$$= \int \mathcal{D}[x(t)] e^{\frac{i}{\hbar} \int dt \mathcal{L}(x(t), \dot{x}(t))}$$

$$= \boxed{\int \mathcal{D}[x(t)] e^{\frac{i}{\hbar} S[x(t)]}}$$

Check formalism: Calculate free particle prop explicitly

Choose $N = 2^A$ for simplicity

$$\text{Calc. } K_N = \left(\frac{mN}{2\pi i \hbar t} \right)^{N/2} \int \prod_{k=1}^{N-1} dx_k e^{i \sum_{k=1}^{N-1} \frac{mN}{2\hbar t} (x_k - x_{k-1})^2}$$

$$\text{Exponent is } \frac{i m N}{2 \hbar t} \left[x_0^2 + 2x_1^2 + 2x_2^2 + \dots + 2x_{N-1}^2 + x_N^2 \right. \\ \left. - 2x_0x_1 - 2x_1x_2 \dots - 2x_{N-1}x_N \right]$$

Do odd integrals first

$$K_N = \prod_{\substack{k=1 \\ k \text{ odd}}}^{N-1} \int dx_k \left[\left(\frac{m 2^A}{2\pi i \hbar t} \right) \int_{(k \text{ odd})} dx_k e^{\frac{i m 2^A}{2 \hbar t} (2x_k^2 + x_{k-1}^2 + x_{k+1}^2 - 2x_k(x_{k-1} + x_{k+1}))} \right]$$

$$\left[2 \left\{ x_k - \frac{1}{2}(x_{k-1} + x_{k+1}) \right\}^2 + \frac{1}{2} x_{k-1}^2 + \frac{1}{2} x_{k+1}^2 - x_{k-1}x_{k+1} \right]$$

$$= \left(\frac{m 2^{A-1}}{2\pi i \hbar t} \right)^{2^{A-2}} \int \prod_{n=1}^{2^{A-1}-1} dx_{2n} e^{\sum_{n=1}^{2^{A-1}-1} \frac{i m 2^{A-1}}{2 \hbar t} (x_{2n} - x_{2n-2})^2}$$

$$= K_{N/2}$$

So by induction,

$$K_N = K_1 = \sqrt{\frac{m}{2\pi i \hbar t}} e^{\frac{im}{2\hbar t} (X_N - X_0)^2}$$

$$= K(X_N, t; X_0, 0) \quad \text{as promised [Exactly].}$$

Feynman path integral approach:

- Alternative formulation of quantum theories.
- Requires classical action $S[X(t)]$ as starting point.
- Requires definition of measure $\mathcal{D}[X(t)]$
- Not practical for most QM calculations
- Highly useful in formulating quantum field theory ("Feynman diagrams")
- Avoids conceptual problems of Hamiltonian formalism of QM.

→ Not a "realist" approach: no $|\psi(t)\rangle$,
can replace by correlation function $\int e^{iS/c} \theta(t_2) \theta(t_1)$

→ Thus, no collapse of wavefunction.

Stationary phase

Given a function $g(x)$, so $\frac{dg}{dx}(x_c) = 0$ at a unique $x = x_c$,

consider $\int dx e^{\frac{i}{\epsilon} g(x)}$, ϵ small.

$$g(x) = g(x_c) + \frac{1}{2} g''(x_c)(x-x_c)^2 + \frac{1}{6} g'''(x_c)(x-x_c)^3 + \dots$$

$$\int dx e^{\frac{i}{\epsilon} g(x)} = e^{\frac{i}{\epsilon} g(x_c)} \sqrt{\frac{2\pi i \epsilon}{g''(x_c)}} \left[1 + O(\epsilon^2) \right]$$

Integral dominated by part near x_c .

Similarly, $\int \mathcal{D}[X(t)] e^{\frac{i}{\hbar} S[X(t)]}$

dominated by X_{class} where $\frac{\delta S}{\delta X}[X_{\text{class}}] = 0$.

$$\text{so } K(x, t; x_0, t_0) \cong e^{iS[X_{\text{class}}(t)]/\hbar}$$

$$\cong e^{iS(x, t; x_0, t_0)/\hbar}$$

For free particle, $S(x, t; x_0, t_0) = \frac{m(x-x_0)^2}{2(t-t_0)}$,

so this is exactly right.

2.4 Quantum particles in potentials and EM fields

Potentials

In classical & Quantum mech, shifting potential by overall constant
 $V \rightarrow V + V_0$
 has no effect on measurable quantities.

Classical: EOM all involve derivatives of V

Newtonian: $F = m - \nabla V$

Hamiltonian: $\{p, V(x)\}$ involves $\partial/\partial x$

Lagrangian: S shifts by $\int V(t) dt$, no effect on δS .

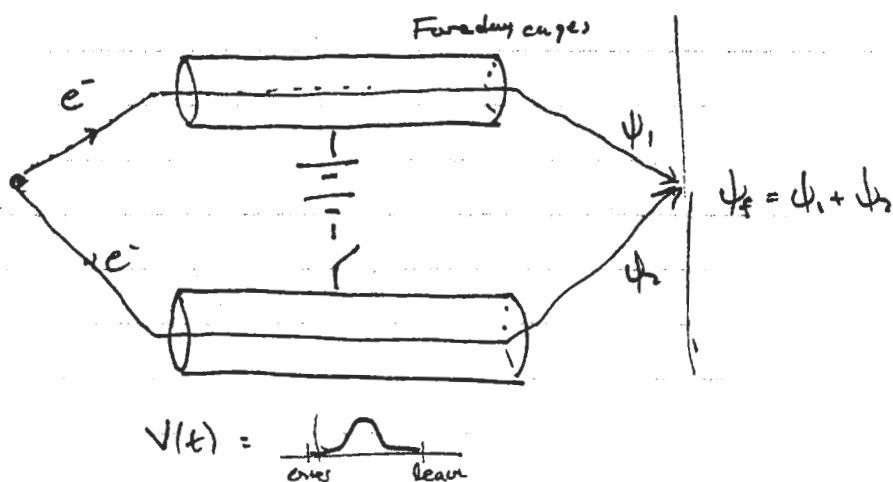
Quantum: $H \rightarrow H + V_0 \Rightarrow$

$$|\psi(t)\rangle \Rightarrow e^{-\frac{i}{\hbar} \int V_0 dt} |\psi(t)\rangle$$

Overall phase ^{is} not observable.

since in expansion $|\psi\rangle = \sum c_a |a\rangle$,
 just changes $c_a \Rightarrow e^{i\theta} c_a$.

Changing potential in one region is observable



without V , by superposition

$$\psi_f = \psi_1 + \psi_2$$

with V ,
$$\psi_f = \left(e^{-\frac{i}{\hbar} \int V(t) dt} \psi_1 + \psi_2 \right)$$

gives phase difference, changes interference pattern.

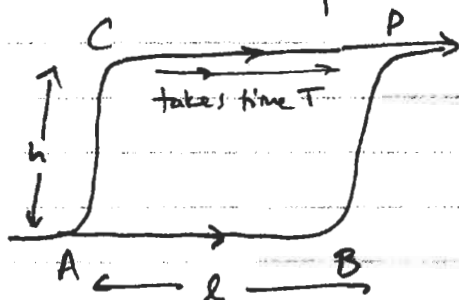
No effect in classical limit $\hbar \rightarrow 0$.

Note: no fields introduced in region with particles (!)
 [this is a variation of Aharonov-Bohm]

Example: Gravitational induced quantum interference

- No quantum theory of gravity
- Hard to see quantum effects where gravity is relevant
 (Gravity $\sim 10^{-39} \times$ as strong as EM forces)
($e^2 N^2 \text{ vs } e^2$)

Possible to see quantum effect through phase difference

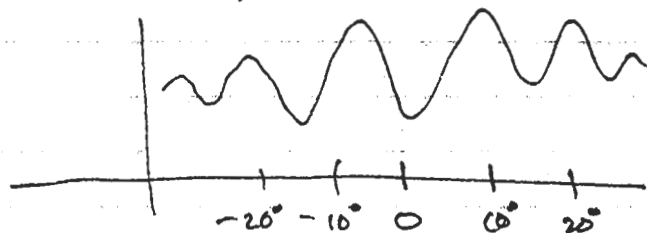


paths ABD vs. ACD:

$$\delta V = mgh$$

Phase difference: $e^{\frac{i}{\hbar} mgh T}$

Interference seen using neutrons following loops rotated @ angle δ from horizontal



Collochia, Overhauser, Werner
1975

Particles in EM fields

Recall Electromagnetism:

Fields $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$

A_μ is 4-vector potential, $A_\mu = (-\phi, \vec{A})$
($x^\mu = (ct, \vec{x})$)

$F_{0i} = -F_{i0} = -E_i$ ($i=1,2,3$)

$F_{ij} = \epsilon_{ijk} B^k$

(Einstein summation convention)

or $E_i = -\frac{1}{c} \frac{\partial \vec{A}}{\partial t} - \frac{\partial \phi}{\partial x^i}$
 $B^i = \epsilon^{ijk} \partial_j A_k$

Gauge invariance: $A_\mu \rightarrow A_\mu + \partial_\mu \Lambda$
leaves $F_{\mu\nu}$, hence E & B unchanged.

Lagrangian for charged particle in EM field is

Relativistic: $S = -m \int d\tau \sqrt{\frac{dx^\mu}{d\tau} \frac{dx_\mu}{d\tau}} + \frac{e}{c} \int d\tau A_\mu \frac{dx^\mu}{d\tau}$
note: δ is total derivative under $\delta A_\mu = \partial_\mu \Lambda$.

Nonrelativistic: $\mathcal{L} = \frac{m}{2} \dot{x}^2 + \frac{e}{c} A_i \dot{x}^i - e\phi$

Going to Hamiltonian formalism:

canonical momentum $p_i = \frac{\partial \mathcal{L}}{\partial \dot{x}^i} = m\dot{x}^i + \frac{e}{c} A^i$

$\mathcal{H} = p_i \dot{x}^i - \mathcal{L}$
 $= \frac{m}{2} \dot{x}^2 + e\phi$

$\mathcal{H} = \frac{1}{2m} \left(\vec{p} - \frac{e}{c} \vec{A} \right)^2 + e\phi$

Classical Poisson bracket: $\{x^i, p_j\} = \delta^i_j$
 (p canonical momentum)

Quantization:

$$x^i \rightarrow \hat{x}^i$$

$$p_j \rightarrow \hat{p}_j = -i\hbar \frac{\partial}{\partial x^j} \quad (\text{note: } p, \text{ not } m\dot{x}, \text{ here})$$

$$H \rightarrow \frac{\hat{p}^2}{2m} - \frac{e}{2mc} (\vec{A} \cdot \hat{p} + \hat{p} \cdot \vec{A}) + \frac{e^2}{2mc^2} A^2 + e\phi.$$

Ehrenfest:

$$m \frac{d^2}{dt^2} \langle \vec{x} \rangle = \frac{d}{dt} \langle \vec{p} \rangle = \frac{1}{i\hbar} \langle [\vec{p}, H] \rangle$$

$$\Rightarrow m \frac{d^2}{dt^2} \langle \vec{x} \rangle = \langle \vec{E}e + \frac{e}{2c} \left(\frac{d\vec{x}}{dt} \times \vec{B} - \vec{B} \times \frac{d\vec{x}}{dt} \right) \rangle$$

(Lorentz force law)

Gauge invariance of QM

Under a gauge transform, $A_\mu \rightarrow A_\mu + \partial_\mu \Lambda$

Classically, canonical momentum $p_i = m\dot{x}^i - \frac{e}{c} A^i$ changes.
 x^i, \dot{x}^i remain fixed.

QM:

Schrödinger $i\hbar \frac{\partial}{\partial t} \psi_{(\vec{x}, t)} = \left[\frac{1}{2m} (-i\hbar \vec{\nabla} - \frac{e}{c} \vec{A})^2 + e\phi \right] \psi_{(\vec{x}, t)}$

Take $\vec{A}' = \vec{A} + \vec{\nabla} \Lambda$
 $\phi' = \phi - \frac{1}{c} \frac{\partial \Lambda}{\partial t}$

Can rewrite Schrödinger

$$\left\{ \left[i\hbar \frac{\partial}{\partial t} - e\phi \right] - \left[\frac{1}{2m} \left(-i\hbar \vec{\nabla} - \frac{e}{c} \vec{A} \right)^2 \right] \right\} \psi(\vec{x}, t) = 0.$$

Clear that $\psi'(\vec{x}, t) = e^{\frac{ie}{\hbar c} \Delta(\phi)} \psi(\vec{x}, t)$

satisfies Schrödinger with $\phi \rightarrow \phi'$, $A \rightarrow A'$:

~~Prove that correct~~

Since

$$\left[i\hbar \frac{\partial}{\partial t} - e\phi' \right] e^{\frac{ie}{\hbar c} \Delta(\phi)} \psi = e^{\frac{ie}{\hbar c} \Delta(\phi)} \left[i\hbar \frac{\partial}{\partial t} - e\phi \right] \psi$$

$$\Delta \left[i\hbar \frac{\partial}{\partial x_i} - \frac{e}{c} A'_i \right] e^{\frac{ie}{\hbar c} \Delta(\phi)} \psi = e^{\frac{ie}{\hbar c} \Delta(\phi)} \left[i\hbar \frac{\partial}{\partial x_i} - \frac{e}{c} A_i \right] \psi.$$

So, under gauge transformations

$$\begin{aligned} \phi &\rightarrow \phi' = \phi - \frac{1}{c} \frac{\partial}{\partial t} \Lambda \\ \vec{A} &\rightarrow \vec{A}' = \vec{A} + \vec{\nabla} \Lambda \\ \psi &\rightarrow \psi' = e^{\frac{ie}{\hbar c} \Delta(\phi)} \psi. \end{aligned}$$

No physical observables change, although, e.g.

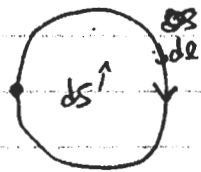
$$\langle \vec{p} \rangle = \langle m\vec{x} + \frac{e}{c} \vec{A} \rangle \text{ is gauge-dependent.}$$

Kinematical momentum $\Pi = m\dot{x}$ is gauge-independent.

Aharonov - Bohm effect

[Another quantum effect arising from fields in regions not containing a particle.]

Recall

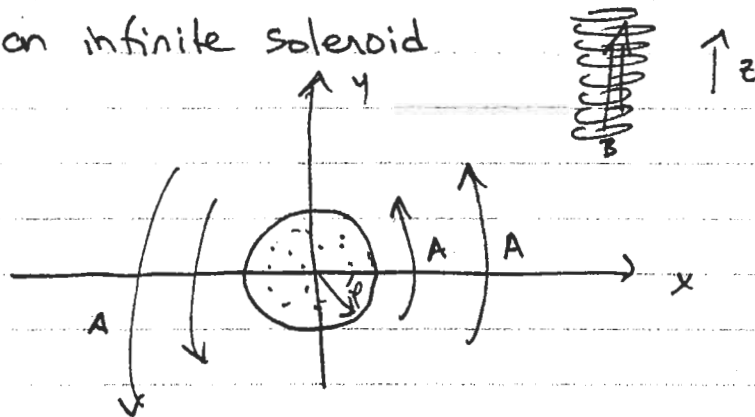


$$\int_{\partial \Sigma} \mathbf{A} \cdot d\mathbf{l} = \int_{\Sigma} (\nabla \times \mathbf{A}) \cdot d\mathbf{s}$$

$$[\text{generally, } \int_{\Sigma} d\omega = \int_{\partial \Sigma} \omega]$$

for differential forms Σ p -manifold
 ω $(p-1)$ -form.

Consider an infinite solenoid

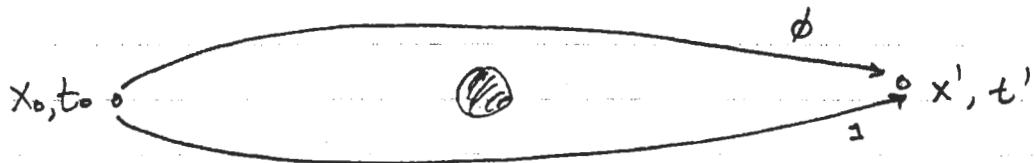


$$\text{Calculate } A: \quad \oint 2\pi R A_{\theta} = \int \mathbf{B} \cdot d\mathbf{s} = \Phi_B$$

$$\text{so } A_{\theta} = \frac{1}{2\pi R} \Phi_B \text{ outside solenoid.}$$

for $R > \rho$

Consider a particle moving around the solenoid
(impenetrable approximation)



Does \mathbf{B} field affect interference pattern? (Yes!)

Use path integrals:

$$K(x_0, t_0; x', t') = \int \mathcal{D}[x(t)] e^{\frac{i}{\hbar} S[x(t)]}$$

Consider paths of type ϕ :

$$e^{\frac{i}{\hbar} S[x(t)]} = e^{\frac{i}{\hbar} \int dt \mathcal{L}(\dot{x}, x)}$$

$$\mathcal{L}(\dot{x}, x) = \frac{m}{2} \dot{x}^2 + \frac{e}{c} A_i \dot{x}^i - e\phi$$

Phase coming from A:

$$e^{\frac{ie}{\hbar c} \int A_i dx^i / dt dt} = e^{\frac{ie}{\hbar c} \int_{\text{path}} A_i dx^i}$$

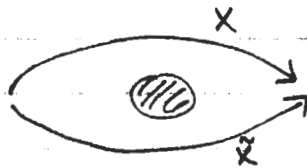
Note that $\int_x A_i dx^i = \int_{\tilde{x}} A_i d\tilde{x}^i$

if $x(t)$, $\tilde{x}(t)$ are topologically equivalent
(i.e., one can be deformed into the other without hitting solenoid)

Thus, all paths of type ϕ give a phase

$$e^{i\theta_0} = e^{\frac{ie}{\hbar c} \int_{\text{path}} A_i dx^i} \quad \text{from A. } \text{Bell}$$

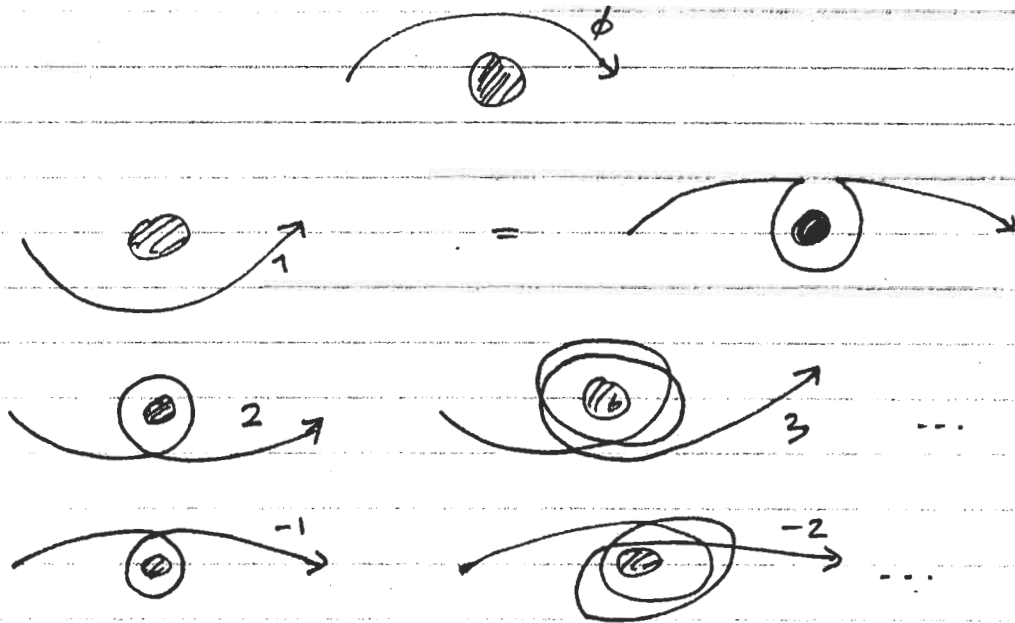
Type 1 paths:



$$\int_{\tilde{x}} A_i d\tilde{x}^i - \int_x A_i dx^i = \Phi_B$$

so $e^{i\theta_0} = e^{i\theta_0 + \frac{ie}{\hbar c} \Phi_B}$

Classify topologically inequivalent paths:

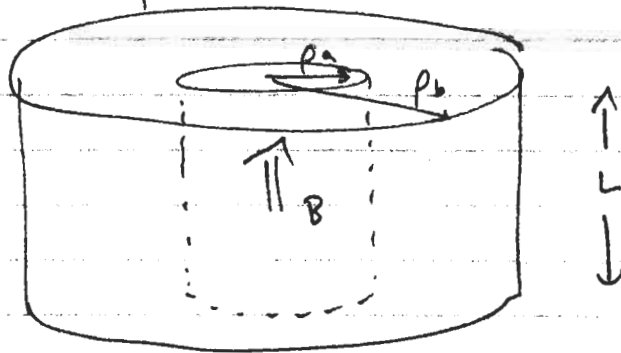


Total propagator:

$$K = \sum_{N=-\infty}^{\infty} \mathcal{D}[x_N(t)] e^{i \frac{e}{\hbar c} \int_0^t A_i dx^i + \frac{i e N}{\hbar c} \Phi_B + \frac{i}{\hbar} \int_{x(t)} \mathcal{L} dx^2}$$

Interference clearly affected by Φ_B .
(paths of type $\phi, 1$ dominate)

Static version of problem



Flux in core, particle in region $r_a < r < r_b$.

Energy levels depend on B (Hw).

Magnetic Monopoles

In a source-free region, Maxwell's equations read:

$$\left. \begin{array}{l} \text{[field notation]} \\ \left\{ \begin{array}{l} \nabla \cdot \vec{E} = 0 \\ \nabla \times \vec{B} = \frac{1}{c} \frac{\partial}{\partial t} \vec{E} \end{array} \right\} \end{array} \right\} \quad \partial_{\mu} F^{\mu\nu} = 0 \quad \left. \begin{array}{l} \text{[form notation]} \\ [d^*F = 0] \end{array} \right\}$$

Since

$$\left. \begin{array}{l} \left\{ \begin{array}{l} \nabla \cdot \vec{B} = 0 \\ \nabla \times \vec{E} = -\frac{1}{c} \frac{\partial}{\partial t} \vec{B} \end{array} \right\} \end{array} \right\} \quad \begin{array}{l} F_{\mu\nu} = \partial_{\mu} A_{\nu} \\ \Downarrow \\ \partial_{\rho} F_{\mu\nu} = \partial_{\rho} \partial_{\mu} A_{\nu} = 0 \end{array} \quad \begin{array}{l} [F = dA] \\ [dF = ddA = 0] \end{array}$$

Equations are invariant under

$$\left\{ \begin{array}{l} \vec{E} \leftrightarrow -\vec{B} \\ \vec{B} \leftrightarrow \vec{E} \end{array} \right\} \quad F_{\mu\nu} \leftrightarrow \frac{1}{2} \epsilon_{\mu\nu\lambda\sigma} F^{\lambda\sigma} \quad [F \leftrightarrow *F]$$

"Maxwell duality"

Including (static) sources:

$$\nabla \cdot \vec{E} = 4\pi\rho.$$

What about magnetic charge $\nabla \cdot \vec{B} = 4\pi\rho_m$?

Note: $\nabla \cdot \vec{B} = 0$ when $\vec{B} = \nabla \times \vec{A}$ $[F = dA]$,

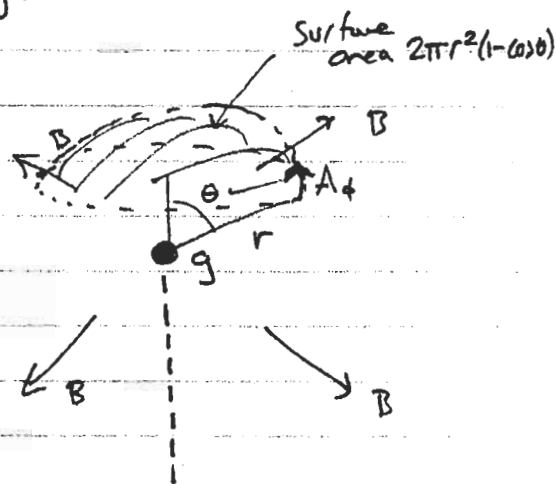
so need to generalize notion of vector potential,
to get ρ_m .

say we have a magnetic charge g , so

$$\vec{B} = \frac{g}{r^2} \hat{r}$$

What is \vec{A} ?

Can try $\vec{A} = A \hat{\phi}$



$$\int \vec{A} \cdot d\vec{l} = 2\pi r \sin\theta A$$

$$= \frac{g}{r^2} \cdot 2\pi r^2 (1 - \cos\theta)$$

[Ex. show $\nabla \times \vec{A} = \vec{B}$ above]

so perhaps
$$\vec{A} = \frac{g(1 - \cos\theta)}{r \sin\theta} \hat{\phi} \quad (?) \text{ [valid for } \theta < \pi \text{]}$$

Singular on z^- axis ("Dirac string")

$$\int dA = 4\pi g$$

Need to use another solution in region outside z^+ :

$$\vec{A} = -\frac{g(1 + \cos\theta)}{r \sin\theta} \hat{\phi} \quad [\theta > 0]$$

\vec{A}, \vec{A} related by a gauge transformation on $0 < \theta < \pi$.

combining local charts \Rightarrow global picture

Mathematically: "Connection on a $U(1)$ fiber bundle over $\mathbb{R}^3 - \{0\}$ with first Chern class 1".

Geometrically: circle over each point in space
connected in topologically nontrivial fashion.

[POV useful for nonabelian gauge theories, Kaluza-Klein theories]

Classically, only \vec{B} is physical, not \vec{A} , so
 "Dirac string" does not pose any obvious problems.

Quantum mechanically,

Recall $e \frac{ie}{\hbar c} \int_P A_\mu dx^\mu$ enters propagator for charged particle moving along path P .

If $\oint A_\mu dx^\mu = 2\pi n \cdot \frac{\hbar c}{e}$, it is not observable.
 Since position is gauge choice, this must be the case.

Thus, we have

$$4\pi g = 2\pi n \cdot \frac{\hbar c}{e},$$

$$\text{so } g = n \cdot \frac{\hbar c}{2|e|} \approx n \left(\frac{137}{2} \right) |e|.$$

Magnetic charge is quantized in units of $\frac{\hbar c}{2|e|}$.

Turning around, assume \exists magnetic monopole of charge g .

Then any electric charge is

$$e = n \frac{\hbar c}{2|g|}.$$

Can explain why proton charge = $|e|$ (known to 4×10^{-9}).

Many models of fundamental physics (GUT's, etc...)
 predict monopoles

No monopoles seen yet in nature [except, possibly, 1].